

# Sampling in Ice Clouds

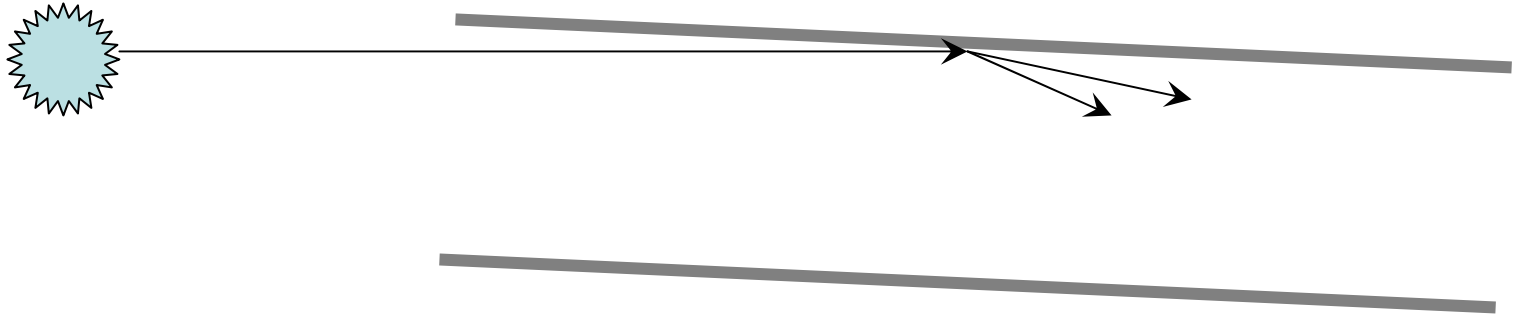
*Daniel Murphy*  
*NOAA Aeronomy Lab*



*also: D. Cziczo, P. Hudson, M. Schein, D. Thomson*

CRYSTAL-FACE Workshop 2003

# The problem:



When it bounces, ice can “sandblast”, shatter, acquire coating, ....

Sticking is unlikely for velocity  $> 1 \text{ m s}^{-1}$ .

It can't completely melt! (insufficient kinetic energy)

Many (10 ??) bounces possible.

# Length and time scales (spherical ice, 190 m s<sup>-1</sup>, 120 mbar)

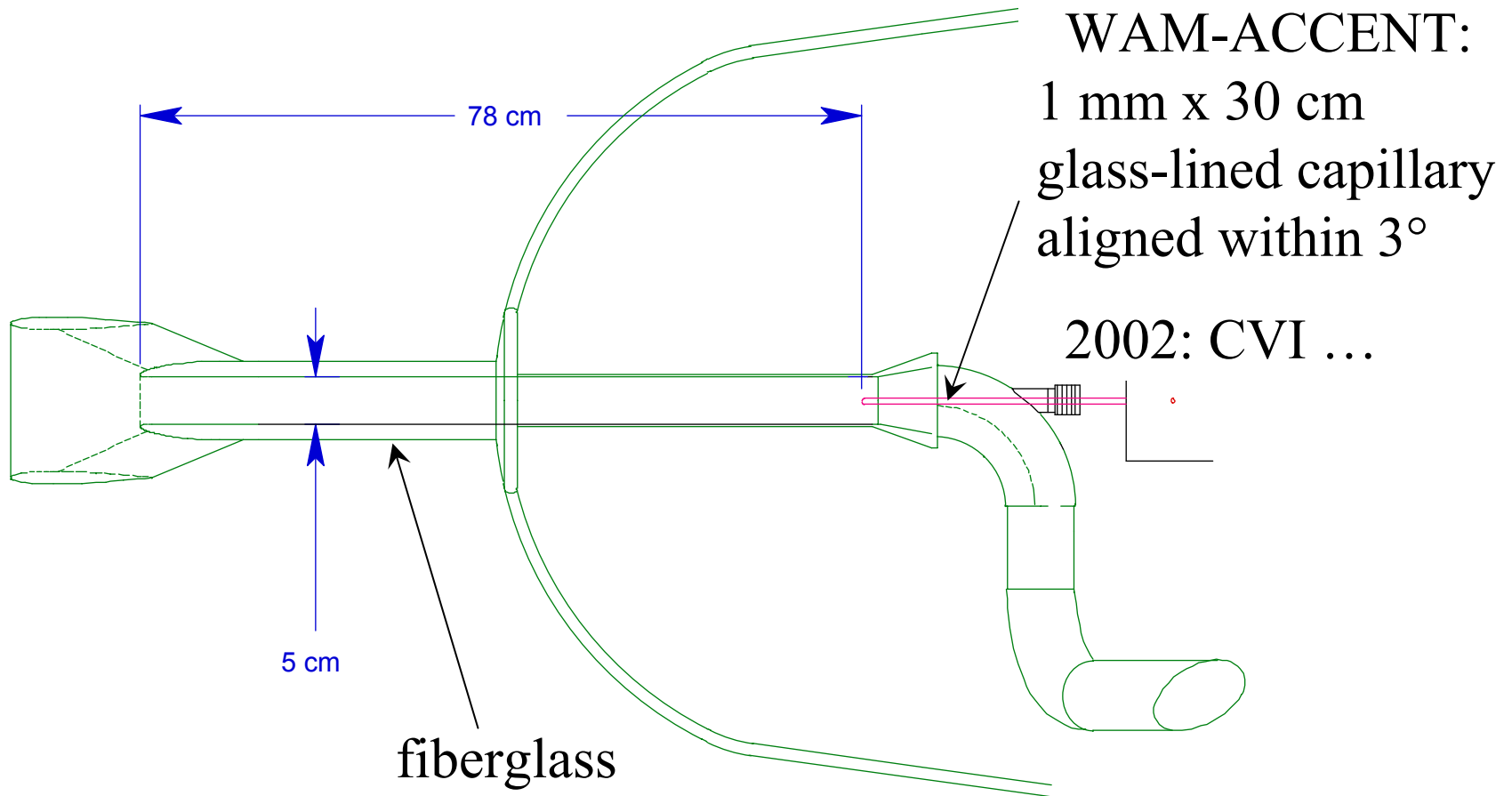
<b>Diameter (<math>\mu\text{m}</math>)</b>	$L_{\text{stop}}$ (cm)	$\tan^{-1}(5\text{mm}/L_{\text{stop}})$ (degrees)	(Grav. fall speed) x (evap. time @ 20°C)
<b>5</b>	1.3	21	25 $\mu\text{m}$
<b>25</b>	14	2	1.3 cm
<b>100</b>	90	0.3	3 m

*$\approx$  inlet alignment requirement.*

*B57 pitch varies by  $\sim 3^\circ$  in “level” flight*

- Large cirrus crystals have moderate Reynolds numbers at aircraft velocity (calculations here include a 1st order correction)
- ( $L_{\text{stop}}$  approx. linear in  $v$ , less than linear in air density)

# PALMS inlet designs

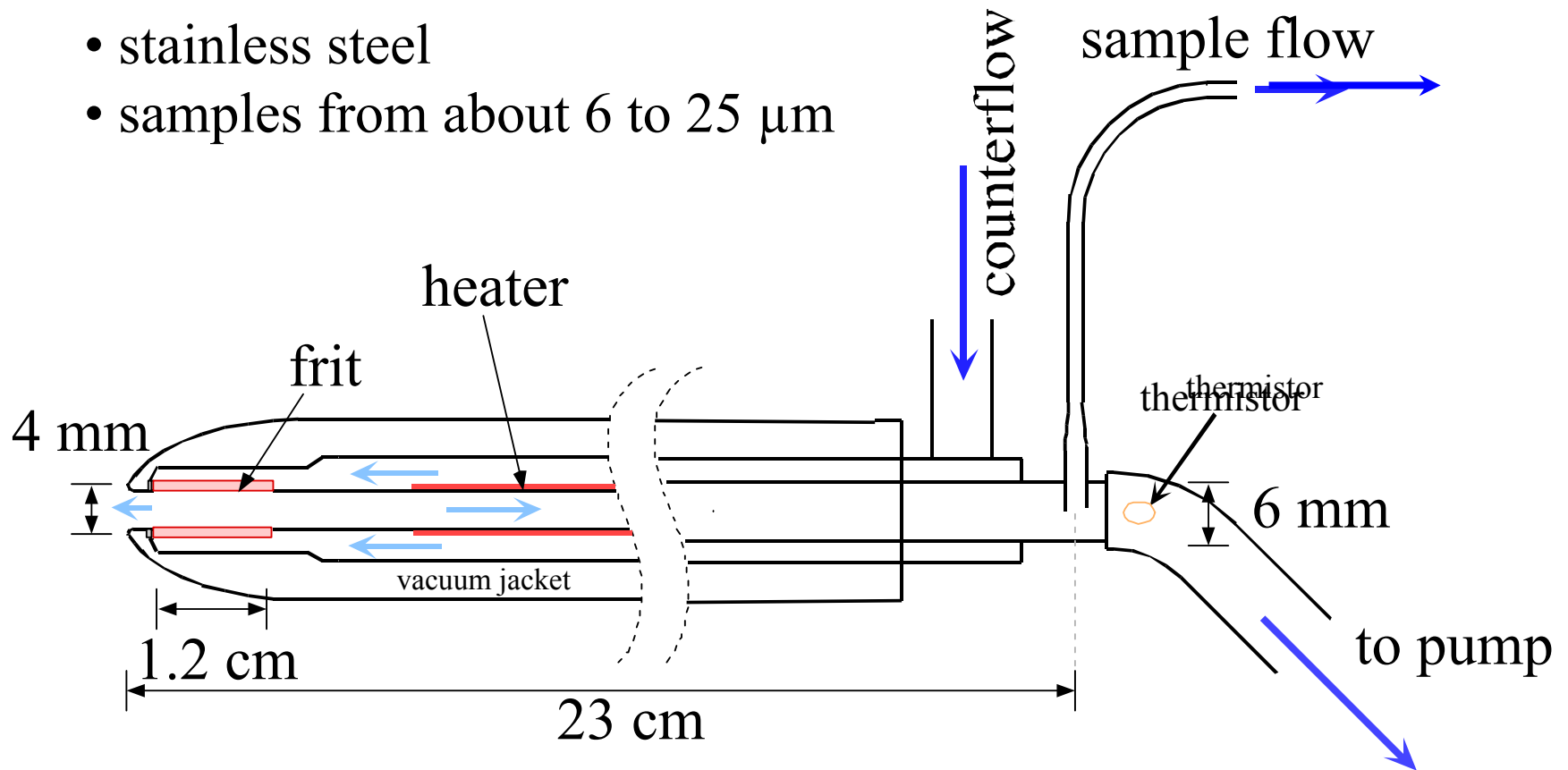


*Even fairly large ice crystals should be aligned with duct*

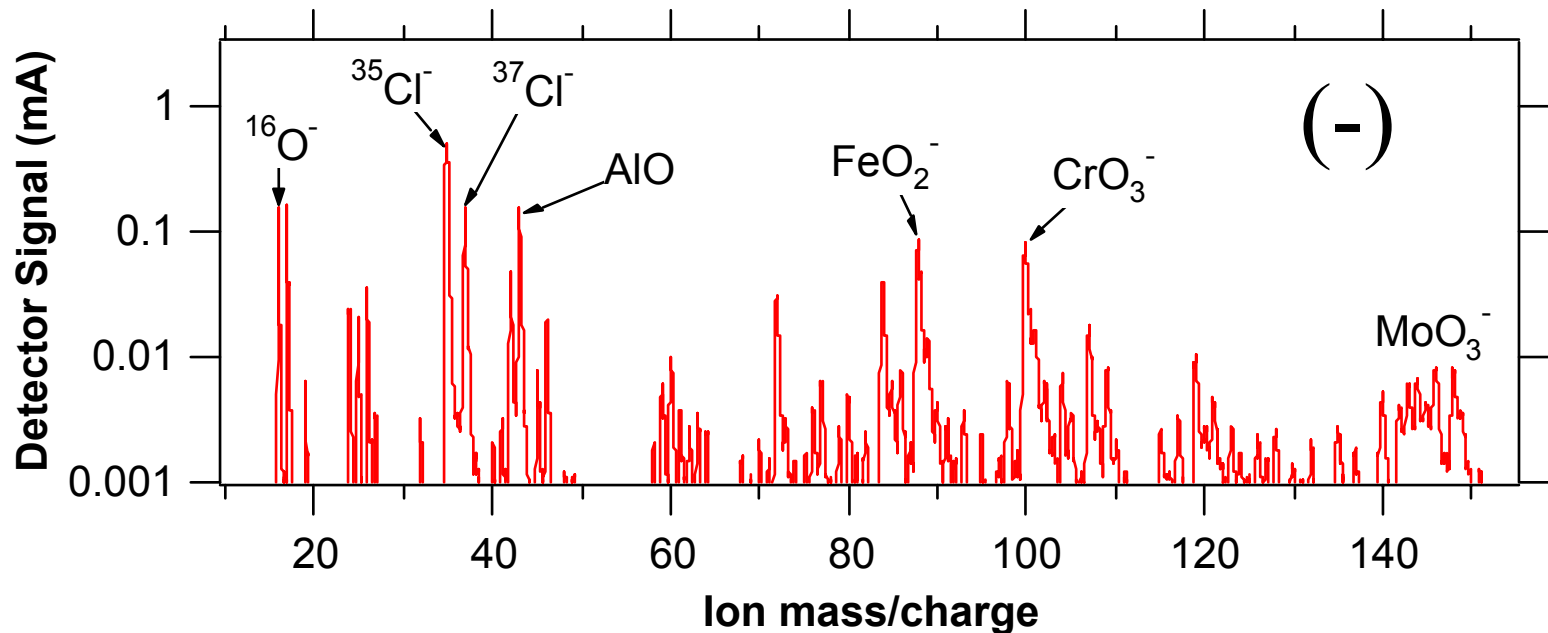
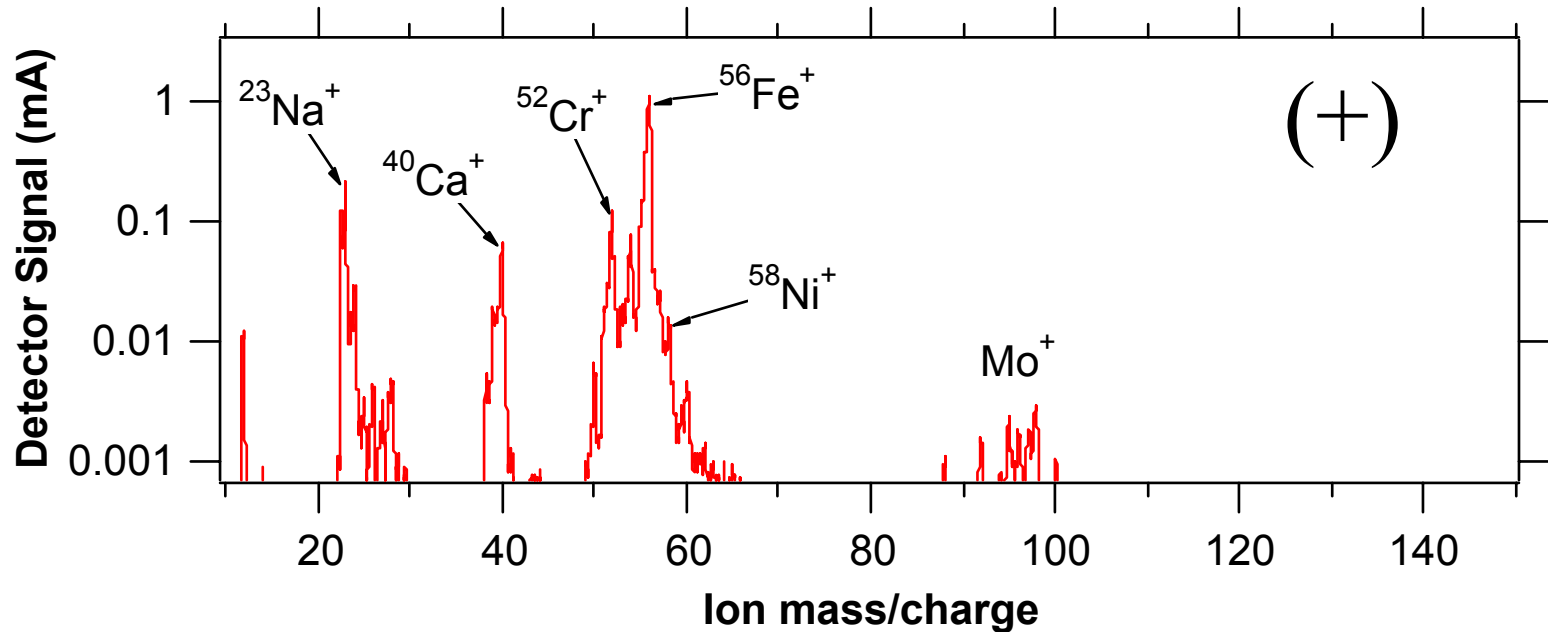
# Counterflow Virtual Impactor (CVI) design

*based on Laucks and Twohy with shorter frit, slightly smaller diameter*

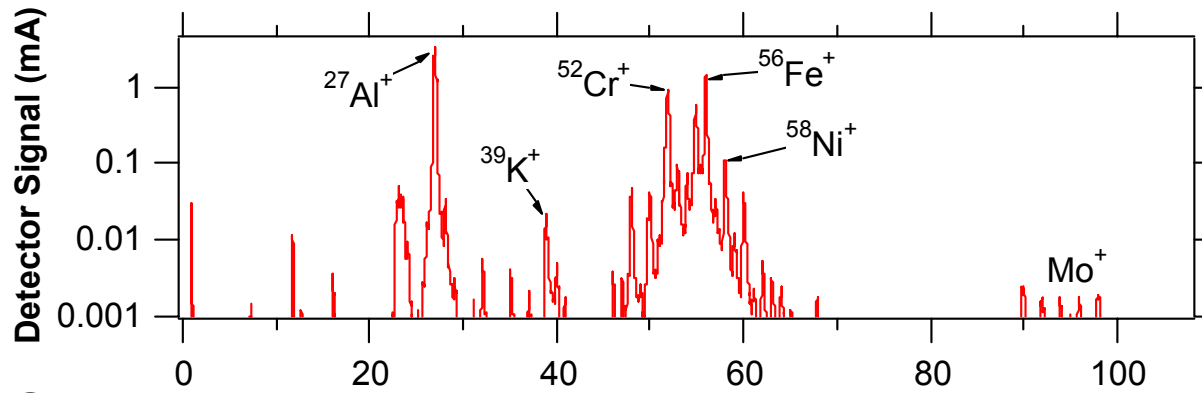
- stainless steel
- samples from about 6 to 25  $\mu\text{m}$



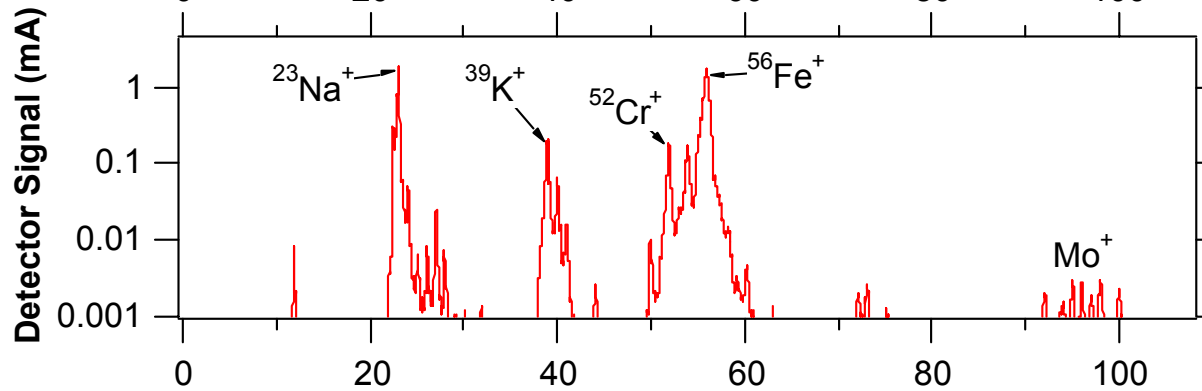
# Typical metal spectra:



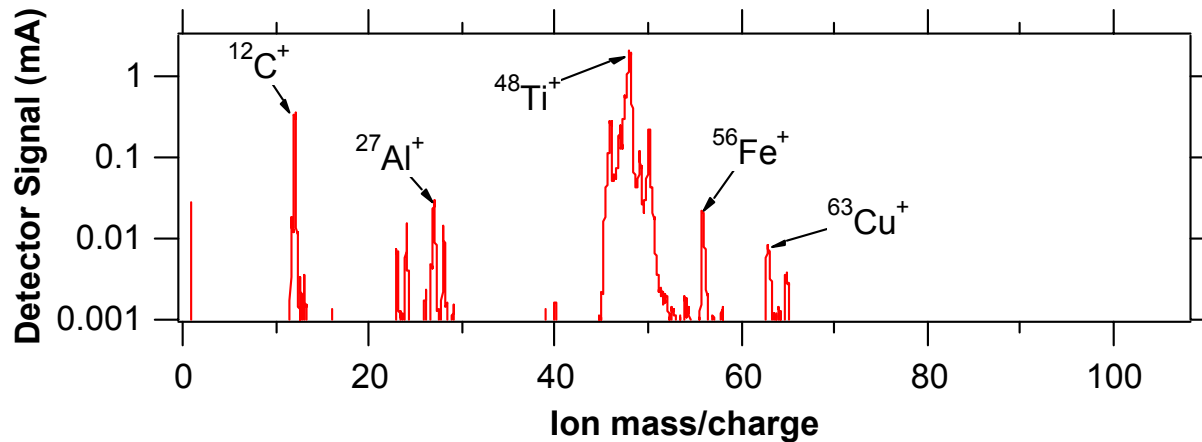
# Variations:



Aluminum



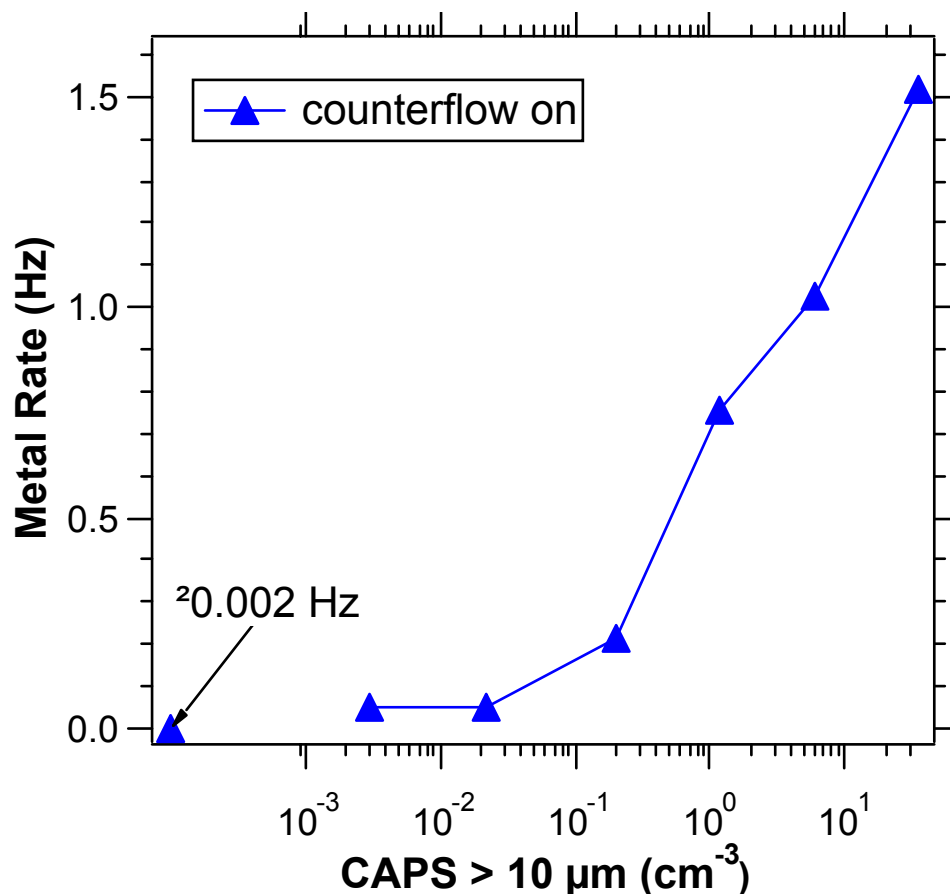
Sodium



Titanium

also Cu, ...

# Frequency of metal particles

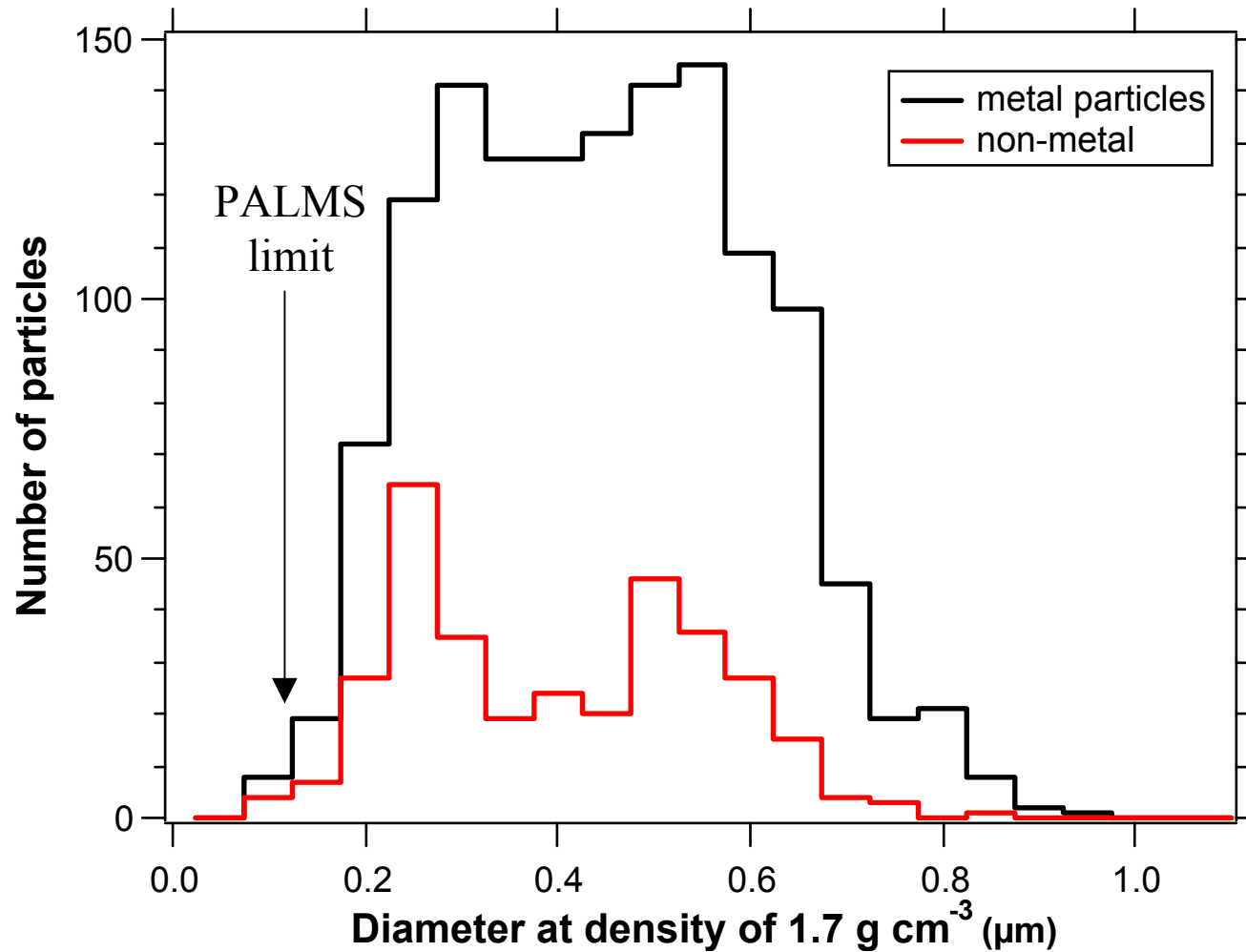


~90% of spectra in  
cloud contained metal

Excludes 709: intermittent CAPS detection of thin cloud, 711-719: PALMS noise affected rates



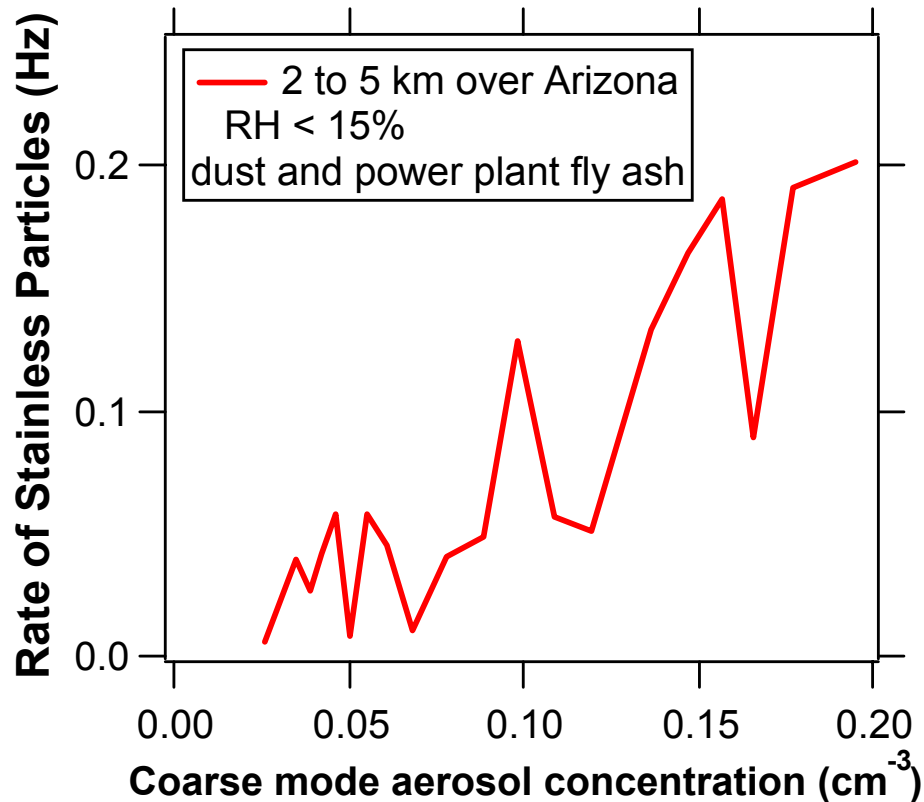
# Size of metal particles



*Uses new  
aerodynamic  
sizing*

*Size is smaller than grain size of steel => uniformity?*

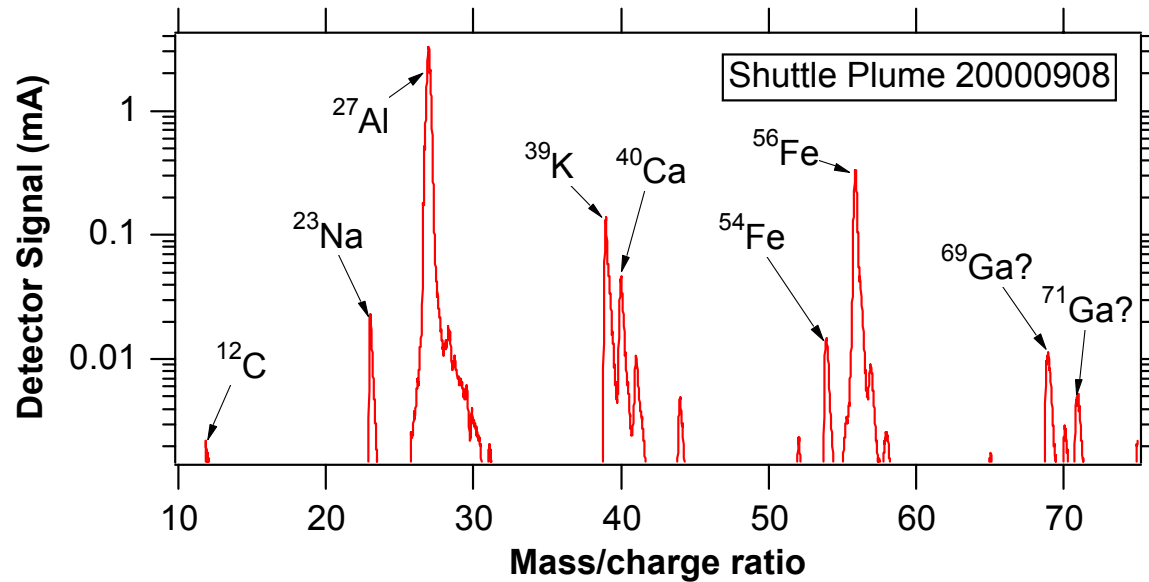
# Dust also creates stainless steel particles



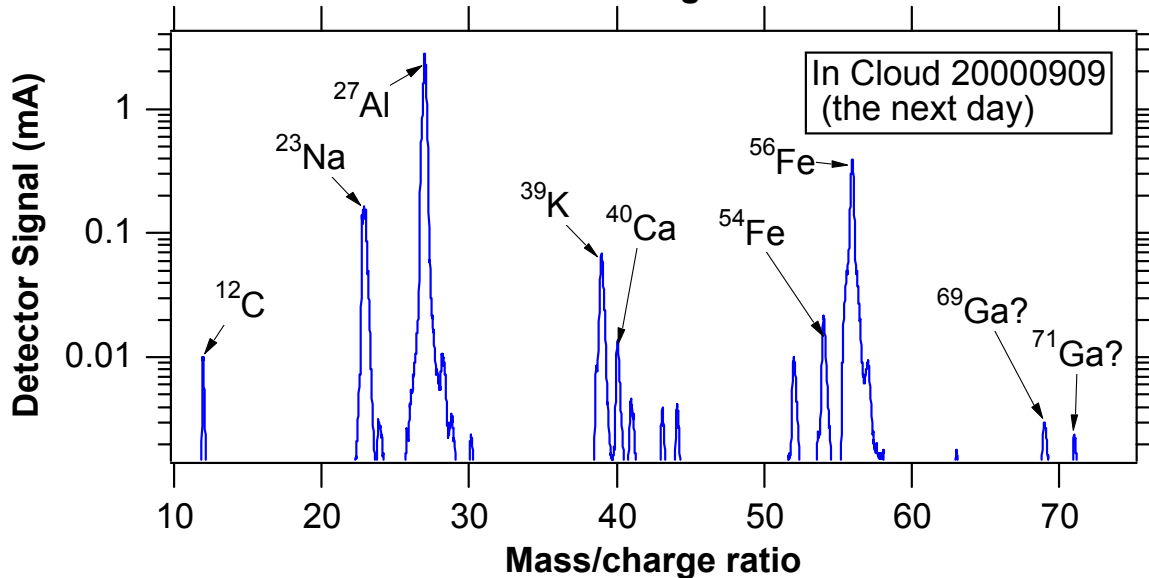
*Mass spectra of these metal particles are similar to metal particles from ice clouds.*

- PALMS flew in a wing pod on the NOAA P3 in spring 2002
- The same inlet hardware as on the B57 in CRYSTAL-FACE (shroud, duct, CVI inlet)

# Previously deposited particles:

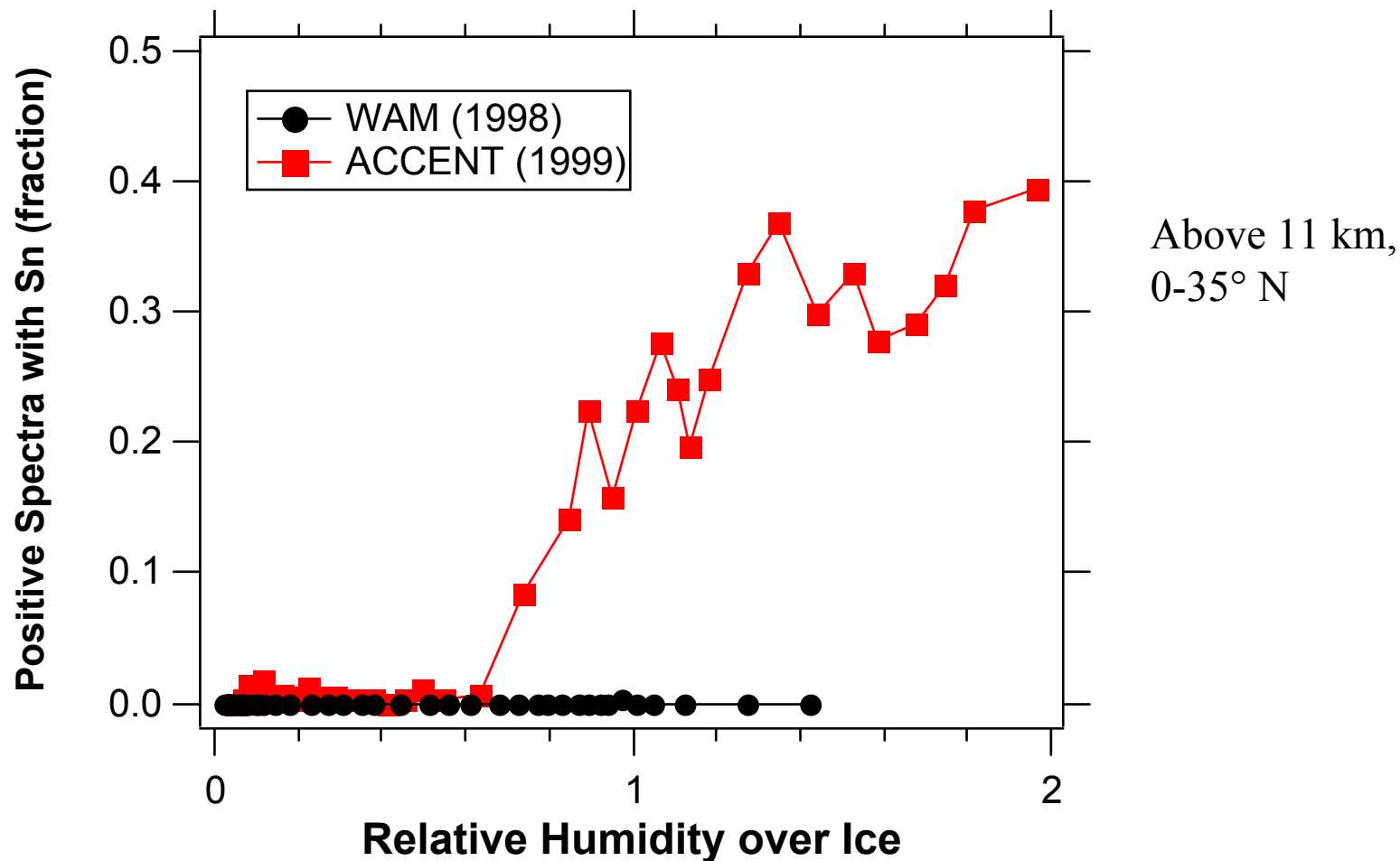


In the space  
shuttle plume



In cloud, the  
next day!

## More wall effects: tin



Hypothesis: Inlet was clean during WAM, contaminated with Sn during ACCENT (by soldering near inlet)

## Previous and future PALMS data?

- *Papers* use data from out of cloud and clean WAM inlet  
=> should be no significant problems

- 
- Some possibly good data will be lost:
    - real metal particles in clouds

- We think we still have good data with uncontaminated particles during CRYSTAL/FACE (-> Cziczo talk)

## Other Groups (*my comments*)

Heintzenberg et al., 1996: CVI, Electron microscope analysis  
cirrus clouds

~75% residues had high Fe content

explained as crustal, *although only 25% also had Si*

Petzold et al., 1998: CVI, Electron microscope analysis

Falcon, both cirrus clouds and contrails

residues  $< 1\ \mu\text{m}$  mostly black carbon

residues  $> 1.5\ \mu\text{m}$  mostly stainless steel

contrail:  $\geq 50\%$  of mass in metals, attributed to engine wear

*ratio to black carbon would imply order 1 ton/yr engine wear*

Noone, K. B., K. H. Noone, J. Heintzenberg, J. Ström, and J. A. Ogren, *J. Atmos. Oceanic Technol.*, **10**, 294, 1993

J. Heintzenberg, K. Okada, and J. Ström, *Atmos. Res.*, **41**, 81, 1996,

Petzold, A., J. Ström, S. Ohlsson, and F. P. Schröder, *Atmos. Res.*, **49**, 21, 1998.

Twohy and Gandrud 1998: CVI, electron microscope  
DC8, contrails

- “Metals” 12 - 36% of residues, typical size 0.4  $\mu\text{m}$
- “Minerals” 30 - 50% of residues, typical size 0.9  $\mu\text{m}$

*Fe particles classified as “mineral”*

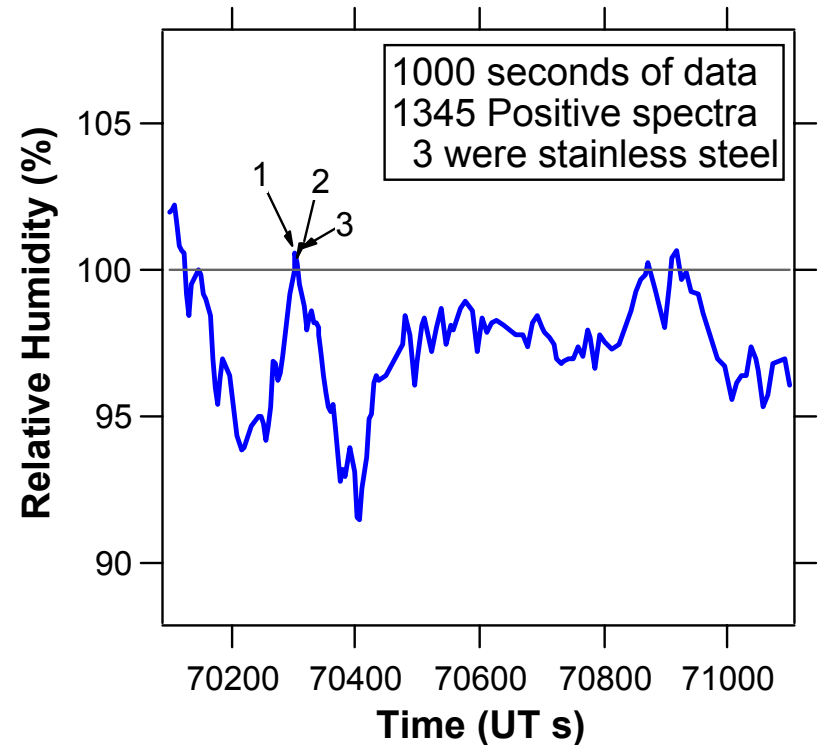
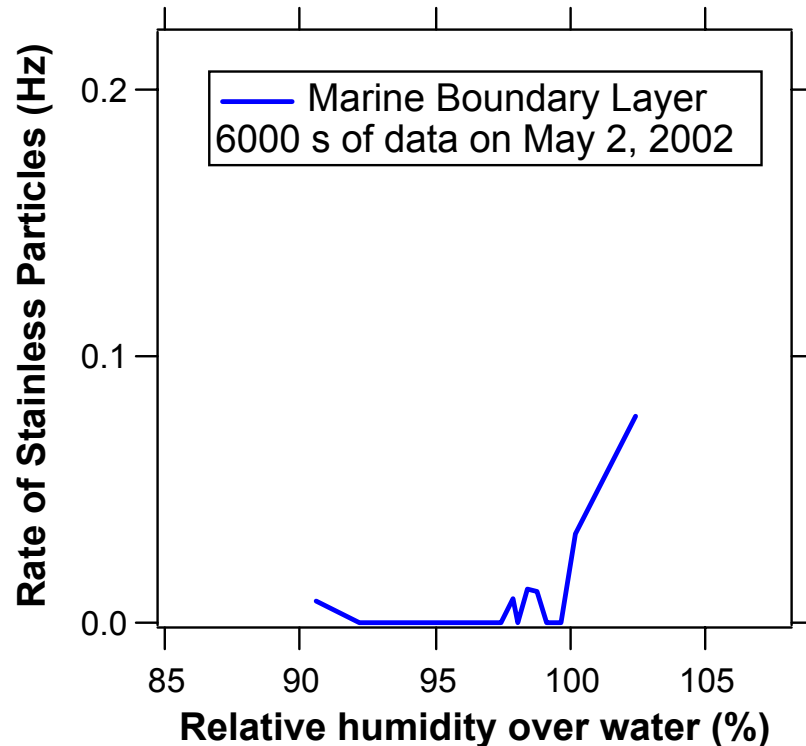
Twohy, C. H., and B. W. Gandrud, *Geophys. Res. Lett.*, **25**, 1359, 1998.

## Other measurements:

- *Everybody* with a CVI has seen metals at significant rates ( $\geq 10\%$ ) but usually less than PALMS (90%).
- Higher rate for PALMS not understood  
PALMS more sensitive to metals?  
Subtle design differences (e.g. shorter frit)?  
Sampling efficiency for different sizes?
- Two CRYSTAL-FACE experiments:  
DU/Arizona State: frequent zinc in ice clouds  
CVI on Citation  
-> See their presentations



# Fast response data are helpful:



*Suppose:* 3 steel particles were found in a 15 minute filter sample near California coast and shipping lanes?

# Could the metal particles be real?

## Pro:

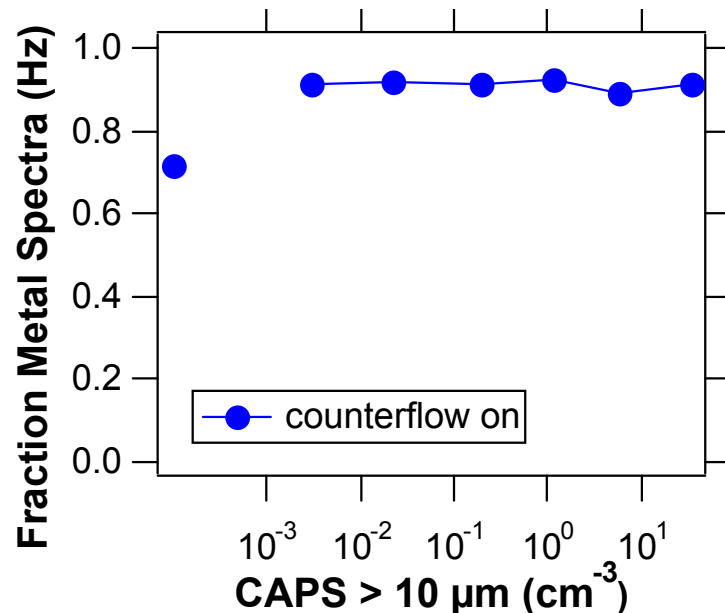
- Sheridan et al. found order 1% metal particles in upper trop.
- Metal particles can be good ice nuclei.

# Could the metal particles be real?

## Con:

- Requires a large source of metals to free troposphere:  
1 per liter, 0.5  $\mu\text{m}$ , Northern Hemisphere, 2 week lifetime  
 $\Rightarrow$  13000 tons/year
- Why stainless steel? (found by both PALMS and Ström et al.)

- Proportion of metal particles stays high even above 1  $\text{cm}^{-3}$  ice



# Are the metal particles inlet contamination?

## Pro:

- Unambiguous correlations of metals with presence of ice or dust.

## Con:

- Metals don't always match inlet composition.

## Notes:

- Chlorine? *Involved in corrosion of stainless steel.*
- Shouldn't the entire front of the airplane be eroded?

*Probably not easy to observe:*

100 l<sup>-1</sup> ice, abrade 0.5 μm chunks => 1 μm in 100 hrs in cirrus

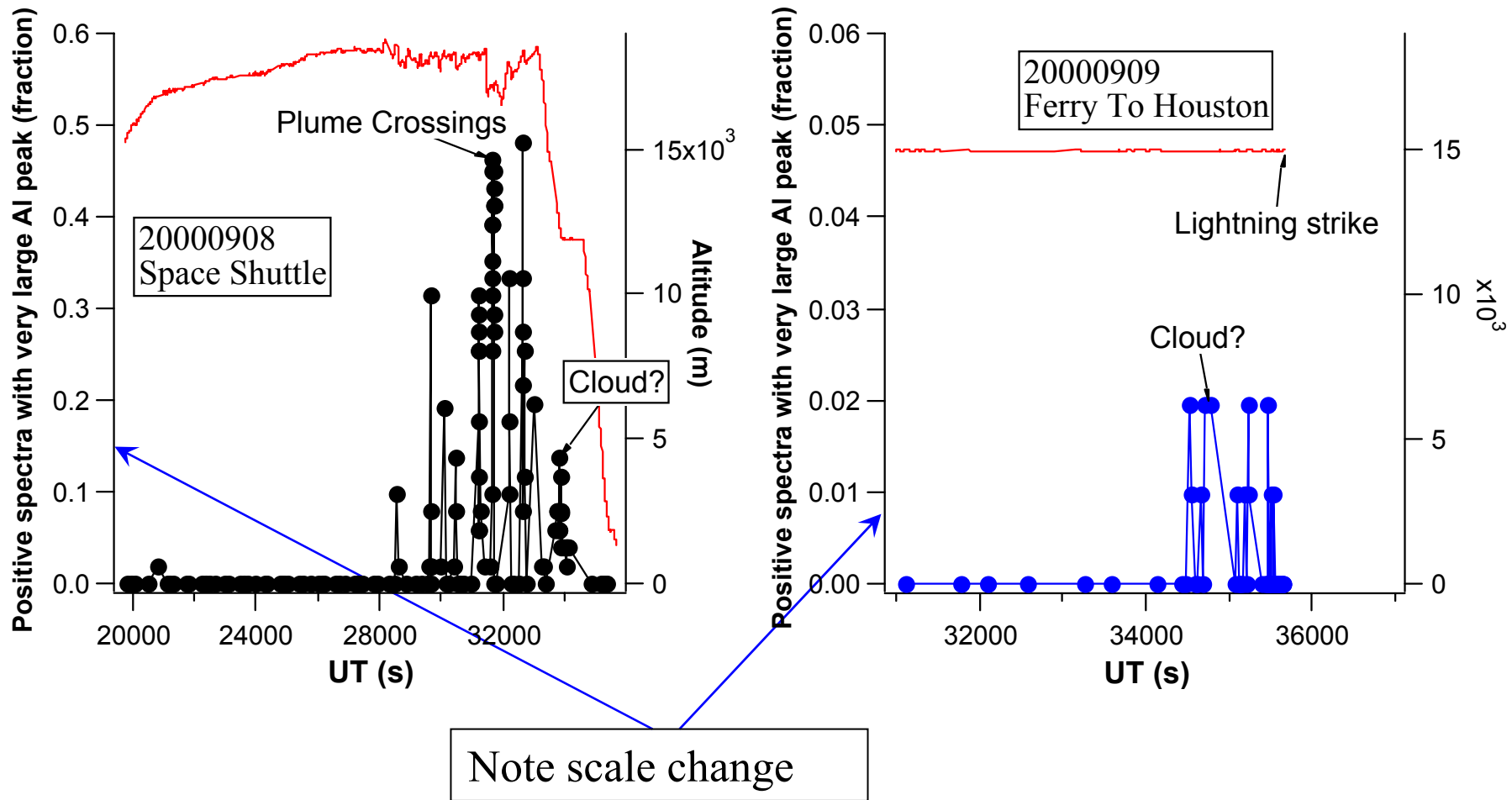
# Summary

- On a high speed aircraft, ice crystals larger than about  $25\text{ }\mu\text{m}$  *will* hit the wall of an inlet, either due to inertia or gravity.
- Published aircraft data on cirrus/contrail ice residues have found substantial numbers ( $>10\%$ ) of metal particles. They cannot be distinguished by size alone.
- Ice crystals can knock pre-existing particles off the wall of an inlet. Ice crystals appear to be able to abrade stainless steel.
- Some real metal particles are possible. However, to explain published CVI data the global flux would be very substantial.
- Abrasion/shattering may be frequent enough to affect data on ice number, especially if knocking older particles off the wall. Water content should be OK.

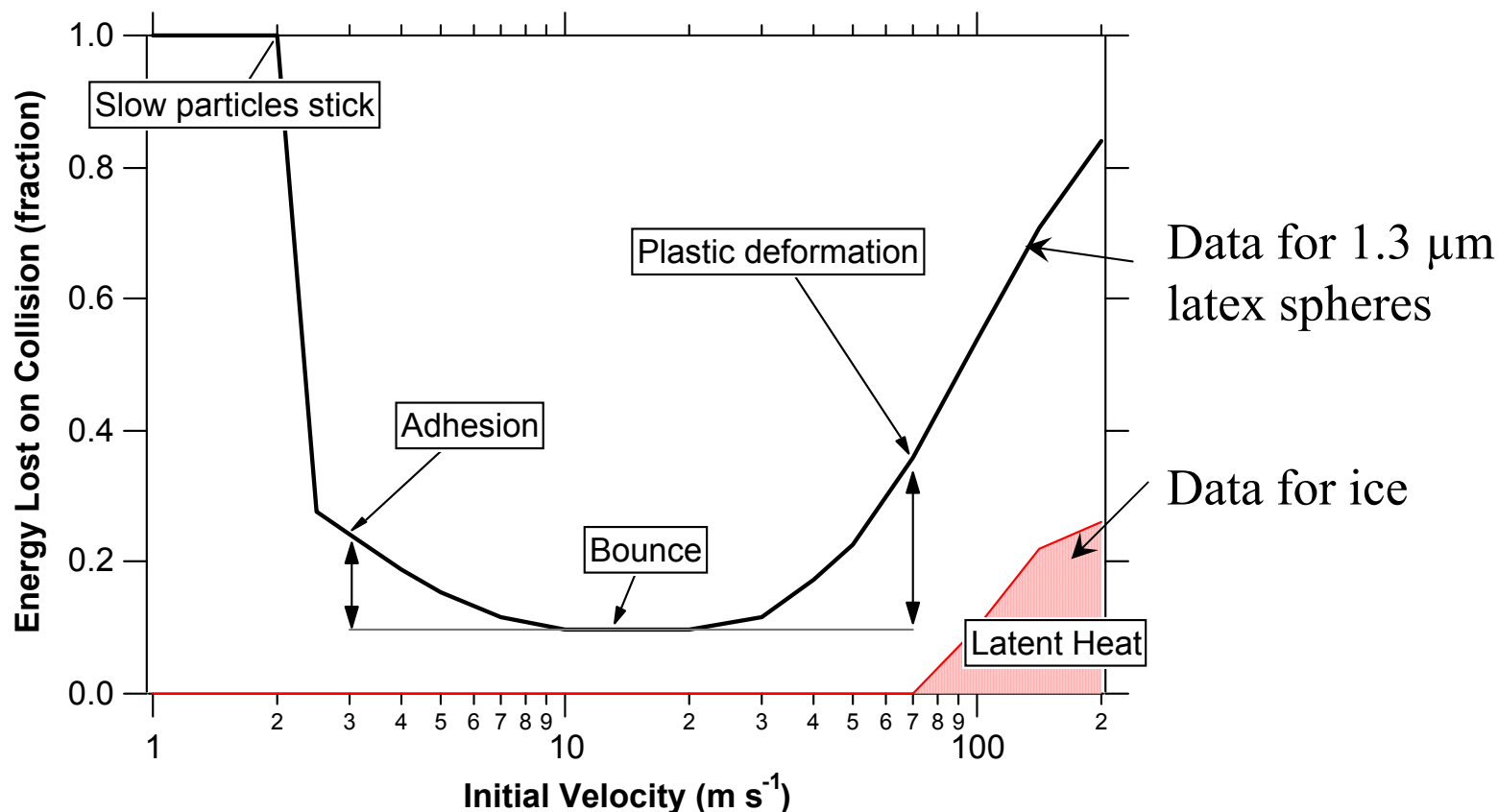
# Suggestions:

- Lab experiments on ice particles hitting steel
- Gold plate the inlet for a unique signature
- Fly in the Southern Hemisphere for less industrial influence.
- Explore “shadowed” forward-facing inlets to exclude the largest particles that always hit a wall.  
(e.g. different version of ER2 football)

# Can ice knock particles off an inlet wall?



# Energetics of wall collisions:



- Larger (e.g. cirrus) particles stick only at  $< 1 \text{ m s}^{-1}$
- Many ( $\sim 10$ ) bounces energetically feasible
- $100 \text{ m s}^{-1}$  collision has more energy than 100 story drop

Dahneke, B., *Aerosol Sci. Technol.*, **23**, 25, 1996.

Sugi, N., M Arakawa, M. Kouchi, and N. Maeno, *Geophys. Res. Lett.*, **25**, 837, 1998.